

ON A TYPE OF $N(k)$ -QUASI EINSTEIN MANIFOLDS

K. LALNUNSIAMI¹ and J.P. SINGH^{*,2}

Abstract

The object of this paper is to study $N(k)$ -quasi Einstein manifolds. W^* -Ricci pseudosymmetric, W_2 -pseudosymmetric and Z -generalized pseudosymmetric $N(k)$ -quasi Einstein manifolds are considered. Finally, we construct examples to prove the existence of such manifolds.

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1 Introduction

Chaki and Maity[4] introduced the notion of a quasi Einstein manifold. A quasi Einstein manifold is a generalization of the Einstein manifold. A non flat Riemannian manifold $(M^n, g)(n > 2)$ is said to be quasi Einstein if the Ricci tensor S is not identically zero and satisfies

$$S(X, Y) = ag(X, Y) + b\eta(X)\eta(Y), \quad (1)$$

$\forall X, Y \in TM$, where a and b are smooth functions, $b \neq 0$ and η is a non-zero 1-form defined by

$$g(X, \xi) = \eta(X), \quad g(\xi, \xi) = \eta(\xi), \quad \eta(\xi) = 1, \quad (2)$$

for the associated vector field ξ . Here, a and b are called the associated scalars and ξ is called the generator of the manifold. Clearly, if $b = 0$, the manifold reduces to an Einstein manifold.

¹Department of Mathematics and Computer Science, Mizoram University, Tanhril, Aizawl 796004, Mizoram, India e-mail: siamiofficial@gmail.com

^{2*} *Corresponding author*, Department of Mathematics and Computer Science, Mizoram University, Tanhril, Aizawl 796004, Mizoram, India e-mail: jpsmaths@gmail.com

Quasi Einstein manifolds has been studied by several authors such as Chaki and Maity[4], De and De[6], De and Ghosh[7], Debnath and Konar[11] and others. The notion of quasi Einstein manifolds has been extended to generalized Einstein manifolds[1], generalized quasi Einstein manifolds([2], [8]), mixed generalized quasi Einstein manifolds[3] and others. Özgür[20] also studied super quasi Einstein manifolds.

In 1988, Tanno[25] defined the k -nullity distribution of a Riemannian manifold as

$$N(k) : p \rightarrow N_p(k) = \{Z \in T_pM : R(X, Y)Z = k[g(Y, Z)X - g(X, Z)Y]\}, \quad (3)$$

$\forall X, Y, Z \in T_pM$, and k is a smooth function. If the generator ξ of a quasi-Einstein manifold belongs to some k -nullity distribution, then it is called an $N(k)$ -quasi Einstein manifold[26]. In an $N(k)$ -quasi Einstein manifold, k is not arbitrary as given by[26]:

Lemma 1. *In an n -dimensional $N(k)$ -quasi Einstein manifold,*

$$k = \frac{a + b}{n - 1}. \quad (4)$$

$N(k)$ -quasi Einstein manifolds have been studied by several authors such as Hui and Lemence[14], Yildiz et al.[27], Singh et al.[22] and others.

In an $N(k)$ -quasi Einstein manifold, we have[26]

$$R(X, Y)\xi = k[\eta(Y)X - \eta(X)Y], \quad (5)$$

$$R(X, \xi)Y = k[\eta(Y)X - g(X, Y)\xi] = -R(\xi, X)Y, \quad (6)$$

$$\eta(R(X, Y)Z) = k[g(Y, Z)\eta(X) - g(X, Z)\eta(Y)]. \quad (7)$$

Pokhariyal and Mishra[21] introduced two types of tensors

$$W_2(X, Y)Z = R(X, Y)Z - \frac{1}{n-1} [g(Y, Z)QX - g(X, Z)QY] \quad (8)$$

and

$$\begin{aligned} W^*(X, Y)Z &= R(X, Y)Z - \frac{1}{2(n-1)} [S(Y, Z)X - S(X, Z)Y \\ &+ g(Y, Z)QX - g(X, Z)QY], \end{aligned} \quad (9)$$

known as the W_2 -curvature tensor and the m -projective curvature tensors respectively, where Q is the Ricci operator.

Mantica and Molinari[17] defined a generalized (0, 2) type tensor known as the Z -tensor as

$$Z(X, Y) = S(X, Y) + \phi g(X, Y), \quad (10)$$

where ϕ is a smooth function. The study of the Z -tensor was continued by Mantica and Molinari[17], Mantica and Suh([18], [19]), etc. In 2016, Mallick and De[16] studied the derivation conditions $R(\xi, X) \cdot Z = 0$ and $P(\xi, X) \cdot Z = 0$ in an $N(k)$ -quasi Einstein manifold, where P is the projective curvature tensor. $N(k)$ -quasi Einstein manifolds satisfying $C(\xi, X) \cdot R = 0$, $R(\xi, X) \cdot W^* = 0$ and $W^*(\xi, X) \cdot S = 0$, where C is the conformal curvature tensor have been studied by De et al.[10]. Also, in 2018, Chaubey[5] studied W^* -pseudosymmetric and Z -recurrent $N(k)$ -quasi Einstein manifolds. These motivated us to study the properties of $N(k)$ -quasi Einstein manifolds.

This paper is organized as follows: After the preliminaries, we study the m -projective curvature tensor in an $N(k)$ -quasi Einstein manifold. In section 4, we consider W^* -Ricci pseudosymmetric $N(k)$ -quasi Einstein manifolds and section 5 deals with W_2 -pseudosymmetric $N(k)$ -quasi Einstein manifolds. Z -generalized pseudosymmetric $N(k)$ -quasi Einstein manifolds are studied in section 6. Finally, we construct examples to support the existence of these manifolds.

2 Preliminaries

Using equations (1) and (2), we obtain

$$S(X, \xi) = (a + b)\eta(X), \tag{11}$$

$$r = na + b, \tag{12}$$

where r is the scalar curvature of the manifold. In an n -dimensional $N(k)$ -quasi Einstein manifold, we have

$$W_2(X, Y)\xi = \frac{b}{(n-1)}[\eta(Y)X - \eta(X)Y], \tag{13}$$

$$W_2(\xi, X)Y = \frac{1}{n-1}[\eta(Y)QX - (a+b)\eta(Y)X], \tag{14}$$

$$\eta(W_2(X, Y)Z) = 0, \tag{15}$$

$$W^*(X, Y)\xi = \frac{b}{2(n-1)}[\eta(Y)X - \eta(X)Y], \tag{16}$$

$$W^*(\xi, X)Y = \frac{b}{2(n-1)}[g(X, Y)\xi - \eta(Y)X], \tag{17}$$

$$\eta(W^*(X, Y)Z) = \frac{b}{2(n-1)}[g(Y, Z)\eta(X) - g(X, Z)\eta(Y)]. \tag{18}$$

The generalized Z -tensor in an $N(k)$ -quasi Einstein manifold takes the form,

$$Z(X, Y) = (a + \phi)g(X, Y) + b\eta(X)\eta(Y), \quad (19)$$

which by contraction, reduces to

$$Z = (a + \phi)n + b. \quad (20)$$

Also,

$$Z(X, \xi) = (a + b + \phi)\eta(X), \quad (21)$$

$$Z(\xi, \xi) = (a + b + \phi), \quad (22)$$

$\forall X, Y, Z \in M^n$.

3 m -projective curvature tensor in an $N(k)$ -quasi Einstein manifold

Suppose an $N(k)$ -quasi Einstein manifold which satisfies

$$W^*(\xi, X).W_2 = 0,$$

or,

$$\begin{aligned} &W^*(\xi, X)W_2(U, V)Z - W_2(W^*(\xi, X)U, V)Z \\ &- W_2(U, W^*(\xi, X)V)Z - W_2(U, V)W^*(\xi, X)Z = 0. \end{aligned} \quad (23)$$

Using (18), (23) it becomes

$$\begin{aligned} &\frac{b}{2(n-1)} \left[g(X, W_2(U, V)Z)\xi - \eta(W_2(U, V)Z)X \right. \\ &- g(X, U)W_2(\xi, V)Z + \eta(U)W_2(X, V)Z \\ &- g(X, V)W_2(U, \xi)Z + \eta(V)W_2(U, X)Z \\ &\left. - g(X, Z)W_2(U, V)\xi + \eta(Z)W_2(U, V)X \right] = 0. \end{aligned} \quad (24)$$

Since $b \neq 0$ and $n > 1$, we have

$$\begin{aligned} &g(X, W_2(U, V)Z)\xi - \eta(W_2(U, V)Z)X \\ &- g(X, U)W_2(\xi, V)Z + \eta(U)W_2(X, V)Z \\ &- g(X, V)W_2(U, \xi)Z + \eta(V)W_2(U, X)Z \\ &- g(X, Z)W_2(U, V)\xi + \eta(Z)W_2(U, V)X = 0. \end{aligned} \quad (25)$$

Taking the inner product of (22) with respect to ξ , we have

$$\begin{aligned} &W_2'(U, V, Z, X) - \eta(W_2(U, V)Z)\eta(X) \\ &- g(X, U)\eta(W_2(\xi, V)Z) + \eta(U)\eta(W_2(X, V)Z) \\ &- g(X, V)\eta(W_2(U, \xi)Z) + \eta(V)\eta(W_2(U, X)Z) \\ &- g(X, Z)\eta(W_2(U, V)\xi) + \eta(Z)\eta(W_2(U, V)X) = 0. \end{aligned} \quad (26)$$

From (16) and (26), it follows that $W'_2(U, V, Z, X) = 0$. Thus, we can state the following theorem:

Theorem 1. *An n -dimensional $N(k)$ -quasi Einstein manifold satisfies the condition $W^*(\xi, X) \cdot W_2 = 0$ if and only if the manifold is W_2 -flat.*

Definition 1. *A Riemannian manifold is said to be semi-symmetric ([23], [24]) if*

$$R \cdot R = 0, \tag{27}$$

where R is the Riemannian curvature tensor.

Consider an $N(k)$ -quasi Einstein manifold which is W^* -semisymmetric. Then, we have

$$(R(X, Y) \cdot W^*)(U, V)Z = 0,$$

which implies that

$$\begin{aligned} R(X, Y)W^*(U, V)Z - W^*(R(X, Y)U, V)Z \\ - W^*(U, R(X, Y)V)Z - W^*(U, V)R(X, Y)Z = 0. \end{aligned} \tag{28}$$

Taking the inner product of (28) with respect to ξ , we have

$$\begin{aligned} g(R(X, Y)W^*(U, V)Z, \xi) - g(W^*(R(X, Y)U, V)Z, \xi) \\ - g(W^*(U, R(X, Y)V)Z, \xi) - g(W^*(U, V)R(X, Y)Z, \xi) = 0. \end{aligned} \tag{29}$$

Substituting $X = \xi$, (29) reduces to

$$\begin{aligned} g(R(\xi, Y)W^*(U, V)Z, \xi) - g(W^*(R(\xi, Y)U, V)Z, \xi) \\ - g(W^*(U, R(\xi, Y)V)Z, \xi) - g(W^*(U, V)R(\xi, Y)Z, \xi) = 0. \end{aligned} \tag{30}$$

Using equations (6) and (17) in (30), we get

$$W^*(U, V, Z, X) - \frac{b}{2(n-1)} [g(U, Y)g(V, Z) - g(V, Y)g(U, Z)] = 0. \tag{31}$$

Making use of (10) and (31), we obtain

$$\begin{aligned} R'(U, V, Z, Y) - \frac{1}{2(n-1)} [S(V, Z)g(U, Y) - S(U, Z)g(V, Y) \\ + S(U, Y)g(V, Z) - S(V, Y)g(U, Z)] \\ - \frac{b}{2(n-1)} [g(U, Y)g(V, Z) - g(V, Y)g(U, Z)] = 0. \end{aligned} \tag{32}$$

Contracting (32) with respect to U and Y , we have

$$S(V, Z) = (a + b)g(V, Z),$$

which is a contradiction as the manifold is quasi Einstein. This leads to the theorem:

Theorem 2. *There does not exist a W^* -semisymmetric $N(k)$ -quasi Einstein manifold.*

Definition 2. *A Riemannian manifold is said to be a symmetric manifold([12], [15]) if*

$$(\nabla_X R)(Y, Z)V = 0, \quad (33)$$

where ∇ is the operator of covariant differentiation with respect to metric g .

Consider an $N(k)$ -quasi Einstein manifold which is W^* -symmetric. Then, we can write

$$(\nabla_X W^*)(U, V, Z, Y) = 0.$$

Using (9), we have

$$\begin{aligned} (\nabla_X R')(U, V, Z, Y) &= \frac{1}{2(n-1)} \left[(\nabla_X S)(V, Z)g(Y, U) - (\nabla_X S)(U, Z)g(V, Y) \right. \\ &\quad \left. + (\nabla_X S)(U, Y)g(V, Z) - (\nabla_X S)(V, Y)g(U, Z) \right]. \end{aligned} \quad (34)$$

Setting $U = Y = e_i$ and summing over $i, 1 \leq i \leq n$, we get

$$(\nabla_X S)(V, Z) = \frac{dr(X)}{n} g(V, Z). \quad (35)$$

Using (2) in (35), we obtain

$$\begin{aligned} da(X)g(V, Z) + db(X)\eta(V)\eta(Z) + b[(\nabla_X \eta)(Z)\eta(V) \\ + (\nabla_X \eta)(Z)\eta(V)] &= \frac{dr(X)}{n} g(V, Z). \end{aligned} \quad (36)$$

Putting $Z = V = \xi$, we get

$$dr(X) = n[da(X) + db(X)]. \quad (37)$$

Also, from (12), it follows that

$$dr(X) = nda(X) + db(X). \quad (38)$$

From (37) and (38), we get

$$db(X) = 0,$$

i.e., b is constant. Therefore, we have

Theorem 3. *There exists no W^* -symmetric $N(k)$ -quasi Einstein manifold unless the associated scalar b is a non-zero constant.*

From (10), we can write

$$\begin{aligned} & (divW^*)(X, Y)Z \\ &= (divR)(X, Y)Z - \frac{1}{2(2n-3)} [dr(X)g(Y, Z) - dr(Y)g(X, Z)], \end{aligned} \quad (39)$$

where “ div ” denotes the divergence.

We know that in a Riemannian manifold,

$$(divR)(X, Y)Z = (\nabla_X S)(Y, Z) - (\nabla_Y S)(X, Z). \quad (40)$$

Using equation (39) in (40), we get

$$\begin{aligned} (divW^*)(X, Y)Z &= (\nabla_X S)(Y, Z) - (\nabla_Y S)(X, Z) \\ &- \frac{1}{2(2n-3)} [dr(X)g(Y, Z) - dr(Y)g(X, Z)]. \end{aligned} \quad (41)$$

Suppose that an $N(k)$ -quasi Einstein manifold is W^* -conservative. Then,

$$(divW^*)(X, Y)Z = 0,$$

or,

$$\begin{aligned} (\nabla_X S)(Y, Z) - (\nabla_Y S)(X, Z) &= \frac{1}{2(2n-3)} [dr(X)g(Y, Z) \\ &- dr(Y)g(X, Z)]. \end{aligned} \quad (42)$$

Making use of (2) in (42), we obtain

$$\begin{aligned} & da(X)g(Y, Z) + db(X)\eta(Y)\eta(Z) - da(Y)g(X, Z) - db(Y)\eta(X)\eta(Z) \\ &+ b[(\nabla_X \eta)(Y)\eta(Z) + (\nabla_X \eta)(Z)\eta(Y) - (\nabla_Y \eta)(X)\eta(Z) - (\nabla_Y \eta)(Z)\eta(X)] \\ &= \frac{1}{2(2n-3)} [dr(X)g(Y, Z) - dr(Y)g(X, Z)]. \end{aligned} \quad (43)$$

Assume that the associated scalar b is non-zero constant. Then $db(X) = 0$, from which it follows that $dr(X) = nda(X)$, $\forall X$. Therefore (43) becomes

$$\begin{aligned} & \frac{3(n-2)}{2(2n-3)} [da(X)g(Y, Z) - da(Y)g(X, Z)] + b[(\nabla_X \eta)(Y)\eta(Z) \\ &+ (\nabla_X \eta)(Z)\eta(Y) - (\nabla_Y \eta)(X)\eta(Z) - (\nabla_Y \eta)(Z)\eta(X)] = 0. \end{aligned} \quad (44)$$

Substituting $Y = Z = \xi$ in (44), we obtain

$$b(\nabla_\xi \eta)(X) = \frac{3(n-2)}{2(2n-3)} [da(X) - da(\xi)\eta(X)]. \quad (45)$$

Contracting (44) over Y and Z , we have

$$b[(\nabla_\xi \eta)(X) + \eta(X) \sum_{i=1}^n (\nabla_{e_i} \eta)(e_i)] - \frac{3(n-1)(n-2)}{2(2n-3)} da(X) = 0. \quad (46)$$

From (45) and (46), it follows that

$$\begin{aligned} b\eta(X) \sum_{i=1}^n (\nabla_{e_i}\eta)(e_i) &= \frac{3(n-1)(n-2)}{2(2n-3)} da(X) \\ &- \frac{3(n-2)}{2(2n-3)} [da(X) - da(\xi)\eta(X)]. \end{aligned} \quad (47)$$

Taking $X = \xi$, (47) becomes

$$\sum_{i=1}^n (\nabla_{e_i}\eta)(e_i) = \frac{3(n-1)(n-2)}{2(2n-3)} da(\xi). \quad (48)$$

Making use of (45) and (48), (46) becomes

$$da(X) = da(\xi)\eta(X). \quad (49)$$

Substituting $X = \xi$ in (44) and using (49), we get

$$b[(\nabla_X\eta)(Y) - (\nabla_Y\eta)(X)] = 0,$$

or,

$$(\nabla_X\eta)(Y) - (\nabla_Y\eta)(X) = 0.$$

which implies that the 1-form η is closed.

Setting $X = \xi$, the above equation reduces to

$$(\nabla_\xi\eta)(Y) = 0.$$

which implies that

$$\nabla_\xi\xi = 0.$$

Therefore, we can state the theorem:

Theorem 4. *On an $(n > 3)$ -dimensional $N(k)$ -quasi Einstein manifold which is W^* -conservative and b is non-zero constant, the associated 1-form η is closed and the integral curves of the generator ξ are geodesics.*

4 W^* -Ricci pseudosymmetric $N(k)$ -quasi Einstein manifold

Definition 3. *A Riemannian manifold is said to be Ricci pseudosymmetric [13] if the tensors $R \cdot S$ and $Q(g, S)$ are linearly dependent at every point of M^n , i. e.,*

$$R \cdot S = L_S Q(g, S),$$

where L_S is a smooth function on $A_S = \{x \in \mathbb{R} : S \neq \frac{r}{n}g \text{ at } x\}$.

Consider an $N(k)$ -quasi Einstein manifold which is W^* -Ricci pseudosymmetric. Then the vectors $W^* \cdot S$ and $Q(g, S)$ are linearly dependent, i.e.,

$$(W^*(X, Y) \cdot S)(Z, U) = L_S Q(g, S)(Z, U; X, Y), \tag{50}$$

where L_S is a function on $A_S = \{x \in \mathbb{R} : S \neq \frac{x}{n}g \text{ at } x\}$. Then,

$$\begin{aligned} S(W^*(X, Y)Z, U) + S(Z, W^*(X, Y)U) &= L_S [S((X \wedge Y)Z, U) \\ &+ S(Z, (X \wedge Y)U)]. \end{aligned} \tag{51}$$

Taking $X = \xi$ in (51), we have

$$\begin{aligned} S(W^*(\xi, Y)Z, U) + S(Z, W^*(\xi, Y)U) &= L_S [S((\xi \wedge Y)Z, U) \\ &+ S(Z, (\xi \wedge Y)U)]. \end{aligned} \tag{52}$$

Using (17) and

$$(X \wedge Y)Z = g(Y, Z)X - g(X, Z)Y, \tag{53}$$

equation (52) becomes

$$\begin{aligned} \left[\frac{b}{2(n-1)} - L_S \right] [S(U, X)g(Y, Z) - S(U, Y)g(X, Z) \\ + g(U, Y)S(Z, X) - g(U, X)S(Z, Y)] = 0, \end{aligned} \tag{54}$$

which implies that either

$$L_S = \frac{b}{2(n-1)},$$

or,

$$\begin{aligned} S(U, X)g(Y, Z) - S(U, Y)g(X, Z) \\ + g(U, Y)S(Z, X) - g(U, X)S(Z, Y) = 0. \end{aligned} \tag{55}$$

Using equation (2), (55) can be written as

$$\begin{aligned} a[g(U, X)g(Y, Z) - g(U, Y)g(X, Z) \\ + g(U, Y)g(Z, X) - g(U, X)g(Z, Y)] \\ + b[\eta(U)\eta(X)g(Y, Z) - \eta(U)\eta(Y)g(X, Z) \\ + g(U, Y)\eta(Z)\eta(X) - g(U, X)\eta(Z)\eta(Y)] = 0. \end{aligned} \tag{56}$$

Contracting equation (56) with respect to X and U , we get

$$g(Y, Z) = n\eta(Y)\eta(Z).$$

Substituting $Y = Z = \xi$ in the above equation, we have

$$n = 1,$$

which is a contradiction. Therefore,

$$L_S = \frac{b}{2(n-1)}.$$

Thus, we can state:

Theorem 5. *An n -dimensional W^* -Ricci pseudosymmetric $N(k)$ -quasi Einstein manifold satisfies the relation $L_S = \frac{b}{2(n-1)}$.*

5 W_2 -pseudosymmetric $N(k)$ -quasi Einstein manifolds

Definition 4. An n -dimensional Riemannian manifold is said to be pseudo-symmetric [13] if

$$R \cdot R = LQ(g, R),$$

i.e., $R \cdot R$ and $Q(g, R)$ are linearly dependent and L is a function on $B = \{x \in \mathbb{R} : \mathbb{R} \neq 0 \text{ at } x\}$.

Suppose that an $N(k)$ -quasi Einstein manifold is W_2 -pseudosymmetric. Then,

$$(R(X, Y) \cdot W_2)(U, V)Z = L_{W_2}Q(g, W_2)(U, V, Z; X, Y), \quad (57)$$

where L_{W_2} is a smooth function on $B_{W_2} = \{x \in \mathbb{R} : W_2 \neq 0 \text{ at } x\}$.

From (57), we have

$$\begin{aligned} & R(X, Y)W_2(U, V)Z - W_2(R(X, Y)U, V) \\ & - W_2(U, R(X, Y)V)Z - W_2(U, V)R(X, Y)Z \\ & = L_{W_2}[(X \wedge_{W_2} Y)W_2(U, V)Z - W_2((X \wedge_{W_2} Y)U, V)Z \\ & - W_2(U, (X \wedge_{W_2} Y)V)Z - W_2(U, V)(X \wedge_{W_2} Y)Z]. \end{aligned} \quad (58)$$

Put $X = \xi$ in the above equation, we have

$$\begin{aligned} & R(\xi, Y)W_2(U, V)Z - W_2(R(\xi, Y)U, V)Z \\ & - W_2(U, R(\xi, Y)V)Z - W_2(U, V)R(\xi, Y)Z \\ & = L_{W_2}[(\xi \wedge_{W_2} Y)W_2(U, V)Z - W_2((\xi \wedge_{W_2} Y)U, V)Z \\ & - W_2(U, (\xi \wedge_{W_2} Y)V)Z - W_2(U, V)(\xi \wedge_{W_2} Y)Z]. \end{aligned} \quad (59)$$

Using (6) and (53), we get

$$\begin{aligned} & (k - L_{W_2})[W_2'(U, V, Z, Y)\xi - \eta(W_2(U, V)Z)Y \\ & - g(Y, U)W_2(\xi, V)Z + \eta(U)W_2(Y, V)Z \\ & - g(Y, V)W_2(U, \xi)Z + \eta(V)W_2(U, Y)Z \\ & - g(Y, Z)W_2(U, V)\xi + \eta(Z)W_2(U, V)X] = 0. \end{aligned} \quad (60)$$

Taking the inner product of (60) with respect to ξ , we get

$$\begin{aligned} & (k - L_{W_2})[W_2'(U, V, Z, Y) - \eta(W_2(U, V)Z)\eta(Y) \\ & - g(Y, U)\eta(W_2(\xi, V)Z) + \eta(U)\eta(W_2(Y, V)Z) \\ & - g(Y, V)\eta(W_2(U, \xi)Z) + \eta(V)\eta(W_2(U, Y)Z) \\ & - g(Y, Z)\eta(W_2(U, V)\xi) + \eta(Z)\eta(W_2(U, V)X)] = 0. \end{aligned} \quad (61)$$

By virtue of (15), (61) reduces to

$$(k - L_{W_2})W_2'(U, V, Z, Y) = 0.$$

Since $W_2 \neq 0$, we have

$$k - L_{W_2} = 0,$$

or,

$$k = L_{W_2}.$$

This leads to the theorem:

Theorem 6. *An $N(k)$ -quasi Einstein manifold is W_2 -pseudosymmetric provided that $k = L_{W_2}$.*

6 Z -generalized pseudosymmetric $N(k)$ -quasi Einstein manifold

Definition 5. *A Riemannian manifold is said to be Ricci-generalized pseudosymmetric [13] if at every point of M^n , the tensors $R \cdot R$ and $Q(S, R)$ are linearly dependent, i. e.,*

$$R \cdot R = LQ(S, R),$$

where L is a function on $A = \{x \in \mathbb{R} : Q(S, R) \neq 0 \text{ at } x\}$.

Consider an $N(k)$ -quasi Einstein manifold which is Z -generalized pseudosymmetric. Then,

$$R \cdot R = L_Z Q(Z, R),$$

where L_Z is a function on $A_Z = \{x \in \mathbb{R} : Q(Z, R) \neq 0 \text{ at } x\}$. Then,

$$\begin{aligned} &R(X, Y)R(U, V)W - R(R(X, Y)U, V)W - R(U, R(X, Y)V)W \\ &- R(U, V)R(X, Y)W = L_Z [(X \wedge_Z Y)R(U, V)W \\ &- R((X \wedge_Z Y)U, V)W - R(U, (X \wedge_Z Y)V)W - R(U, V)(X \wedge_Z Y)W]. \end{aligned} \quad (62)$$

Taking $X = \xi$ in (62), we have

$$\begin{aligned} &R(\xi, Y)R(U, V)W - R(R(\xi, Y)U, V)W - R(U, R(\xi, Y)V)W \\ &- R(U, V)R(\xi, Y)W = L_Z [(\xi \wedge_Z Y)R(U, V)W \\ &- R((\xi \wedge_Z Y)U, V)W - R(U, (\xi \wedge_Z Y)V)W - R(U, V)(\xi \wedge_Z Y)W]. \end{aligned} \quad (63)$$

Using (6) and

$$(X \wedge_Z Y)U = Z(Y, U)X - Z(X, U)Y,$$

in (63), we have

$$\begin{aligned}
& [k - L_Z(a + \phi)] [R'(U, V, W, Y)\xi - \eta(R(U, V)W)Y \\
& -g(Y, U)R(\xi, V)W + \eta(U)R(Y, V)W \\
& -g(Y, V)R(U, \xi)W + \eta(V)R(U, Y)W \\
& -g(Y, W)R(U, V)\xi + \eta(W)R(U, V)Y] \\
& = L_Z b [\eta(Y)\eta(R(U, V)W)\xi - \eta(R(U, V)W)Y \\
& -g(Y, U)R(\xi, V)W + \eta(U)R(Y, V)W \\
& -g(Y, V)R(U, \xi)W + \eta(V)R(U, Y)W \\
& -g(Y, W)R(U, V)\xi + \eta(W)R(U, V)Y]. \tag{64}
\end{aligned}$$

Taking the inner product of (64) with respect to ξ , we have

$$\begin{aligned}
& [k - L_Z(a + \phi)] [R'(U, V, W, Y) - \eta(R(U, V)W)\eta(Y) \\
& -g(Y, U)\eta(R(\xi, V)W) + \eta(U)\eta(R(Y, V)W) \\
& -g(Y, V)\eta(R(U, \xi)W) + \eta(V)\eta(R(U, Y)W) \\
& -g(Y, W)\eta(R(U, V)\xi) + \eta(W)\eta(R(U, V)Y)] \\
& = L_Z b [\eta(Y)\eta(R(U, V)W) - \eta(R(U, V)W)\eta(Y) \\
& -g(Y, U)\eta(R(\xi, V)W) + \eta(U)\eta(R(Y, V)W) \\
& -g(Y, V)\eta(R(U, \xi)W) + \eta(V)\eta(R(U, Y)W) \\
& -g(Y, W)\eta(R(U, V)\xi) + \eta(W)\eta(R(U, V)Y)]. \tag{65}
\end{aligned}$$

Using (3) and (8), (65) reduces to

$$L_Z b k [\eta(W)\eta(U)g(V, Y) - \eta(W)\eta(V)g(U, Y)] = 0,$$

which implies (since $b \neq 0$),

$$L_Z k = 0,$$

i.e., $L_Z = 0$ or $k = 0$.

This leads to the theorem:

Theorem 7. *A Z -generalized pseudosymmetric $N(k)$ -quasi Einstein manifold is either semisymmetric or $k = 0$.*

7 Examples of $N(k)$ -quasi Einstein manifolds

Example 1: Consider a Riemannian metric g on \mathbb{R}^3 by

$$ds^2 = g_{ij} dx^i dx^j = e^{x^3} \cos(x^3) [(dx^1)^2 + (dx^2)^2] - (dx^3)^2.$$

Then, we have

$$\begin{aligned}
g_{11} &= g_{22} = e^{x^3} \cos(x^3), & g_{33} &= -1, \\
g^{11} &= g^{22} = e^{-x^3} \sec(x^3), & g^{33} &= -1.
\end{aligned}$$

Then, the non-vanishing components of the Christoffels symbols and the curvature tensors are

$$\Gamma_{11}^3 = \Gamma_{22}^3 = e^{x^3} \frac{\cos(x^3) - \sin(x^3)}{2},$$

$$\Gamma_{13}^1 = \Gamma_{23}^2 = \frac{\cos(x^3) - \sin(x^3)}{2\cos(x^3)},$$

$$R_{1221} = -e^{2x^3} \frac{(1 - \sin(2x^3))}{4}, \quad R_{1331} = R_{2332} = -e^{x^3} \frac{(1 + \sin(2x^3))}{4\cos(x^3)}.$$

Also, the non-vanishing components of the Ricci tensors are

$$S_{11} = S_{22} = -e^{x^3} \sin(x^3), \quad S_{33} = \frac{(1 + \sin(2x^3))}{2\cos^2(x^3)}.$$

Using these results in

$$r = g^{ij} S_{ij}, \tag{66}$$

we get

$$r = -\frac{(\sec^2(x^3) - 6\tan^2(x^3))}{2},$$

which is non-zero.

To show that the manifold is $N(k)$ -quasi Einstein, we choose the scalar functions a and b and the 1-form η as

$$a = -\tan(x^3), \quad b = \frac{1}{2}\sec^2(x^3),$$

$$\eta_i(x) = \begin{cases} 1, & i = 3, \\ 0, & \text{otherwise,} \end{cases}$$

at any point $x \in \mathbb{R}^3$.

From (1), we have

$$S_{11} = ag_{11} + b\eta_1\eta_1, \tag{67}$$

$$S_{22} = ag_{22} + b\eta_2\eta_2, \tag{68}$$

$$S_{33} = ag_{33} + b\eta_3\eta_3 \tag{69}$$

and all others hold trivially.

$$\begin{aligned}
\text{R. H. S of (69)} &= ag_{33} + b\eta_3\eta_3 \\
&= -\tan(x^3)(-1) + \frac{1}{2}\sec^2(x^3)(1) \\
&= \frac{(1 + \sin(2x^3))}{2\cos^2(x^3)} = S_{33} \\
&= \text{L.H.S of (69)}.
\end{aligned}$$

Similarly, it can be shown that equations (67) and (68) hold. Using (66), we get

$$k = \frac{a + b}{n - 1} = \frac{\sin(2x^3) - 1}{4}.$$

So, (\mathbb{R}^3, g) is an $N(\frac{\sin(2x^3)-1}{4})$ -quasi Einstein manifold.

Example 2: Consider \mathbb{R}^4 with the Riemannian metric g defined by

$$ds^2 = g_{ij}dx^i dx^j = (x^3)^2[(dx^1)^2 + (dx^2)^2 + (dx^3)^2] + (dx^4)^2.$$

Then, we have

$$\begin{aligned}
g_{11} = g_{22} = g_{33} &= (x^3)^2, & g_{44} &= 1, \\
g^{11} = g^{22} = g^{33} &= \frac{1}{(x^3)^2}, & g^{44} &= 1.
\end{aligned}$$

The non-vanishing components of the Christoffels symbols, the curvature tensors and the Ricci tensors are

$$\Gamma_{11}^3 = \Gamma_{22}^3 = -\frac{1}{x^3}, \quad \Gamma_{33}^3 = \Gamma_{13}^1 = \Gamma_{23}^2 = \frac{1}{x^3},$$

$$R_{1331} = R_{2332} = -1, \quad R_{2332} = 1, \quad S_{11} = S_{22} = S_{44} = 0, \quad S_{33} = \frac{2}{(x^3)^2}.$$

Using (66) and the above results, we get

$$r = 2,$$

which is non-vanishing. To show that the manifold under consideration is an $N(k)$ -quasi Einstein manifold, we choose the scalar functions a, b and the 1-form η as

$$a = 0, \quad b = 2,$$

$$\eta_i(x) = \begin{cases} \frac{1}{x^3}, & i = 3, \\ 0, & \text{otherwise,} \end{cases}$$

at any point $x \in \mathbb{R}^4$. From (2), we have

$$S_{11} = ag_{11} + b\eta_1\eta_1, \tag{70}$$

$$S_{22} = ag_{22} + b\eta_2\eta_2, \quad (71)$$

$$S_{33} = ag_{33} + b\eta_3\eta_3, \quad (72)$$

$$S_{44} = ag_{44} + b\eta_4\eta_4 \quad (73)$$

and all others hold trivially.

$$\begin{aligned} \text{R. H. S of (72)} &= ag_{33} + b\eta_3\eta_3 \\ &= -0 + 2\left(\frac{1}{x^3}\right)\left(\frac{1}{x^3}\right) \\ &= \frac{2}{(x^3)^2} = S_{33} \\ &= \text{L.H.S of (72)}, \end{aligned}$$

Similarly, it can be shown that equations (70), (71) and (72) hold. Using (4), we get

$$k = \frac{a+b}{n-1} = \frac{2}{3}.$$

So, (\mathbb{R}^4, g) is an $N\left(\frac{2}{3}\right)$ -quasi Einstein manifold.

Example 3: A perfect fluid pseudo Ricci-symmetric spacetime is an $N\left(\frac{2r}{9}\right)$ -quasi Einstein manifold[9].

Example 4: A four dimensional conformally flat perfect fluid (M^4, g) is an $N\left(\frac{1}{3}\left(\frac{1}{2}r + f(T) + e\pi\rho^2(p + \rho)f'(T)\right)\right)$ -quasi Einstein manifold[5].

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